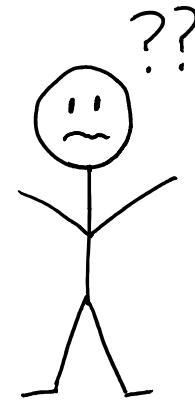
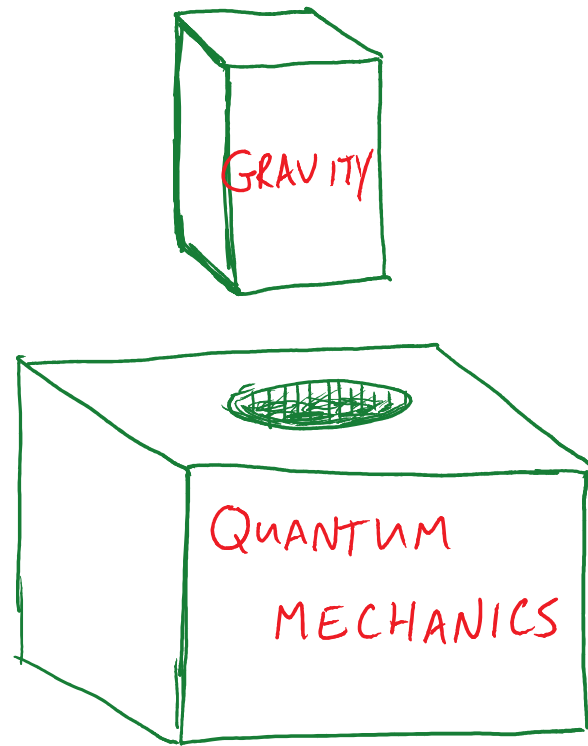


# GRAVITY AND ENTANGLEMENT

Mark Van Raamsdonk, UBC

TESTING GRAVITY 2015

One of the greatest challenges for  
theoretical physics:



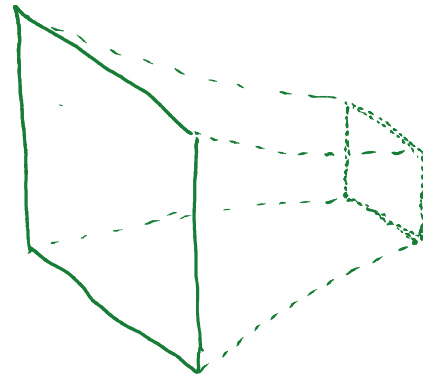
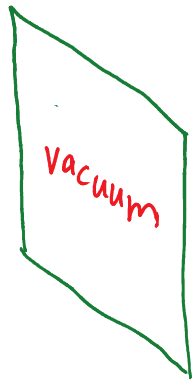
1997: Remarkable progress via AdS/CFT  
correspondence in string theory (Maldacena)

QUANTUM  
GRAVITY  
(certain  
examples)

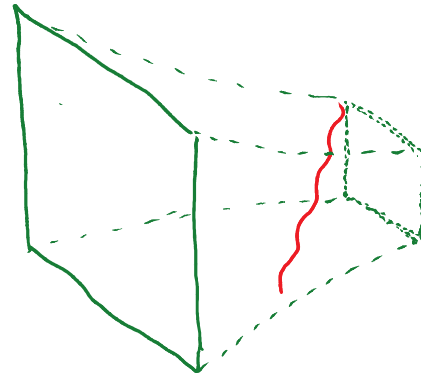
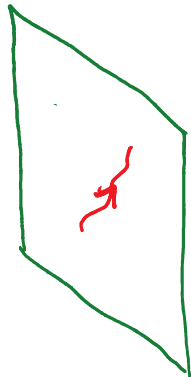
=  
exactly  
equivalent

ORDINARY  
QUANTUM  
SYSTEM  
(eg. QFT on  
fixed spacetime)

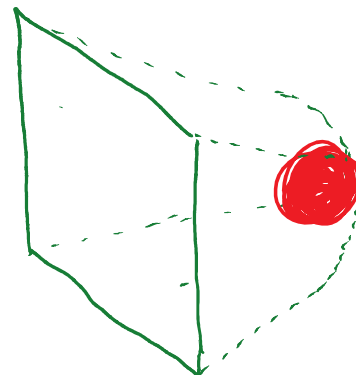
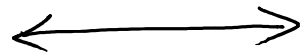
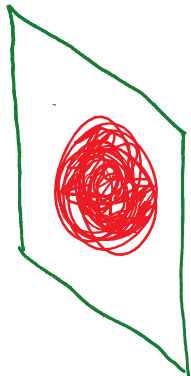
Different QFT states  $\longleftrightarrow$  Different spacetimes



empty  
spacetime



gravity  
wave



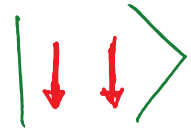
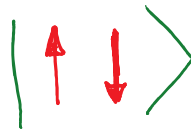
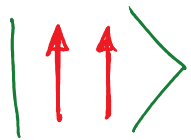
black  
hole

BIG QUESTION:

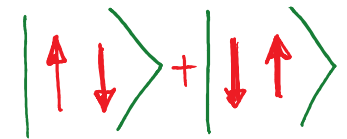
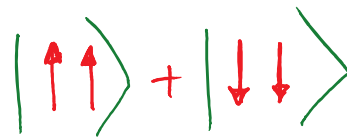
How/why do spacetime/gravity  
emerge from QFT physics?

Recent work: Quantum entanglement is crucial.

e.g. 2 spin system



Not entangled



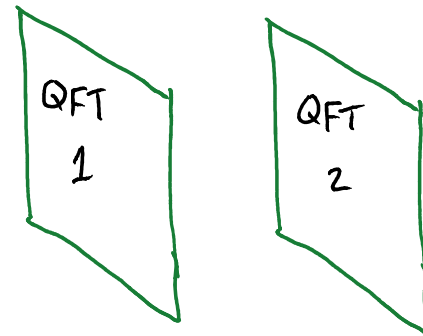
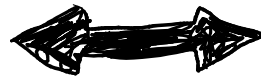
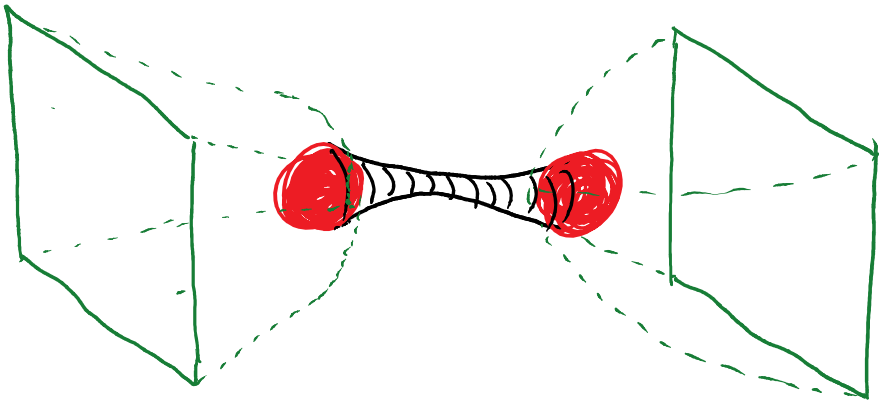
Entangled

Maldacena 2001:

2 separate spacetimes  
connected by wormhole

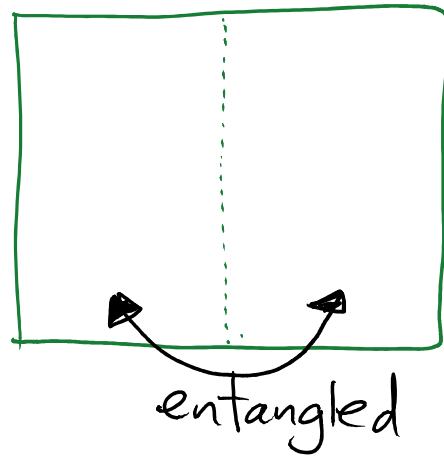


2 separate QFTs  
entangled with  
one another



$$\begin{aligned} |\psi\rangle &= | \text{empty} \text{ empty} \rangle + | \text{particle} \text{ particle} \rangle \\ &+ | \text{particle} \text{ hole} \rangle + | \text{hole} \text{ particle} \rangle \\ &+ \dots \end{aligned}$$

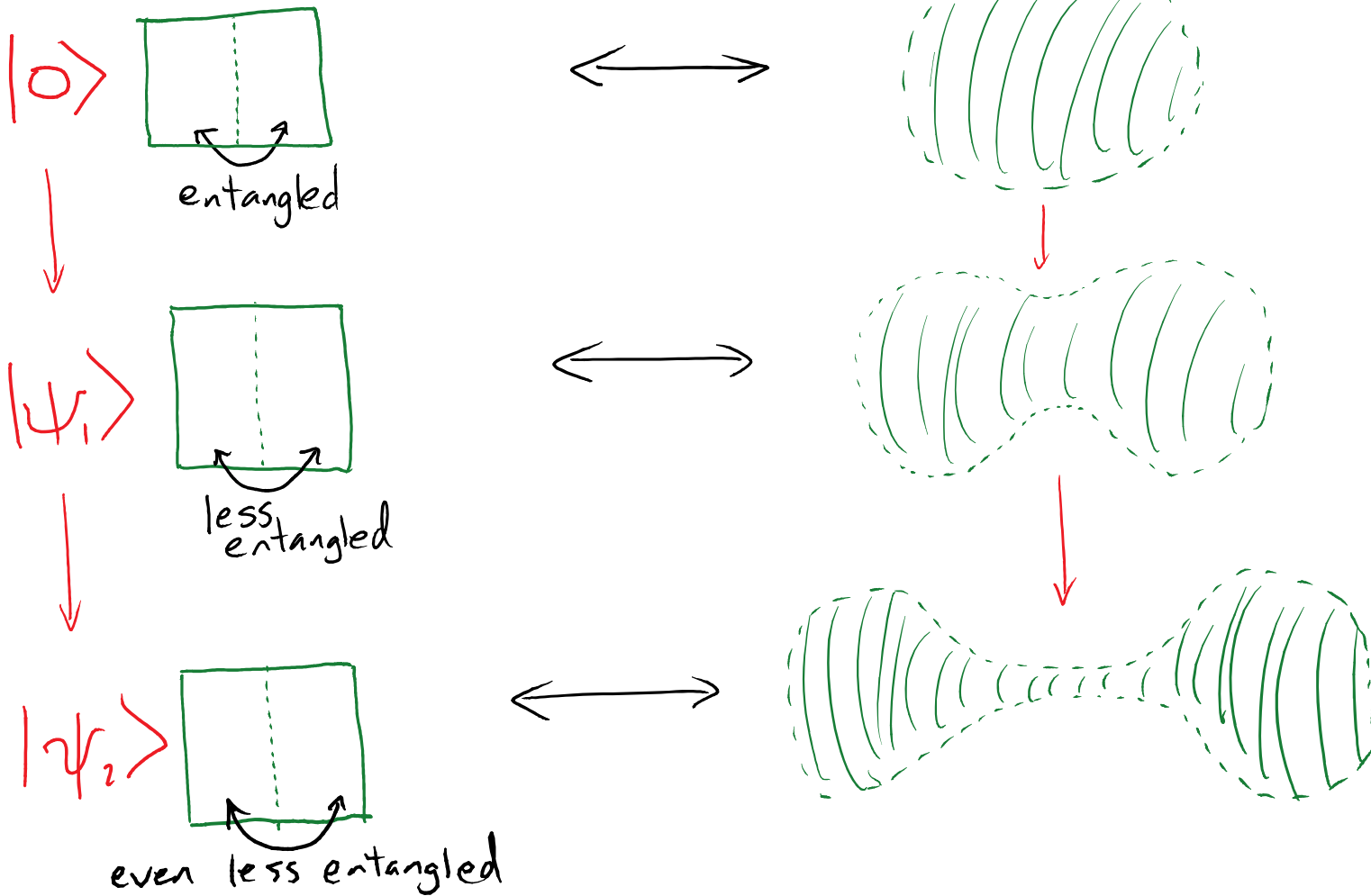
BUT: lots of entanglement already in QFT ground state dual to empty spacetime.



QFT fields in any region entangled with fields outside



What happens if we remove this entanglement?



2 regions of space pinch off from each other!

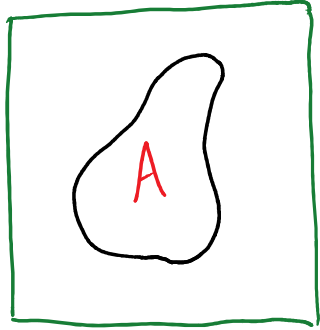
MIR; WR, Czech, Noguiera, Karzmarek

Suggests that classical spacetime geometry  
emerges via entanglement of degrees of freedom  
in dual QFT!

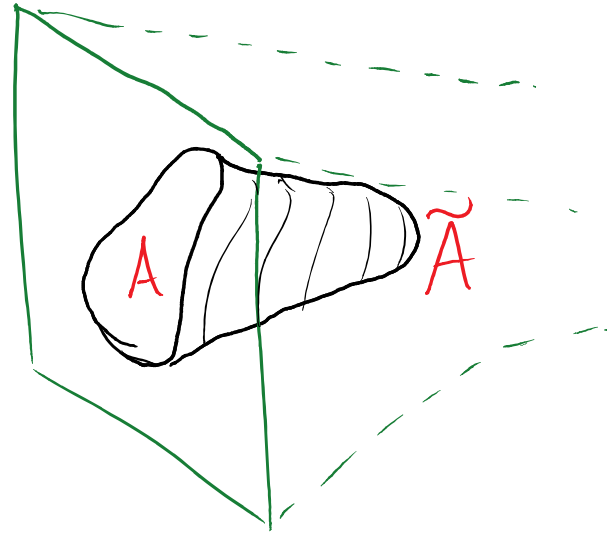
MVR, Swingle 2009

No classical spacetime without quantum entanglement.

# A quantitative connection (Ryu + Takayanagi)

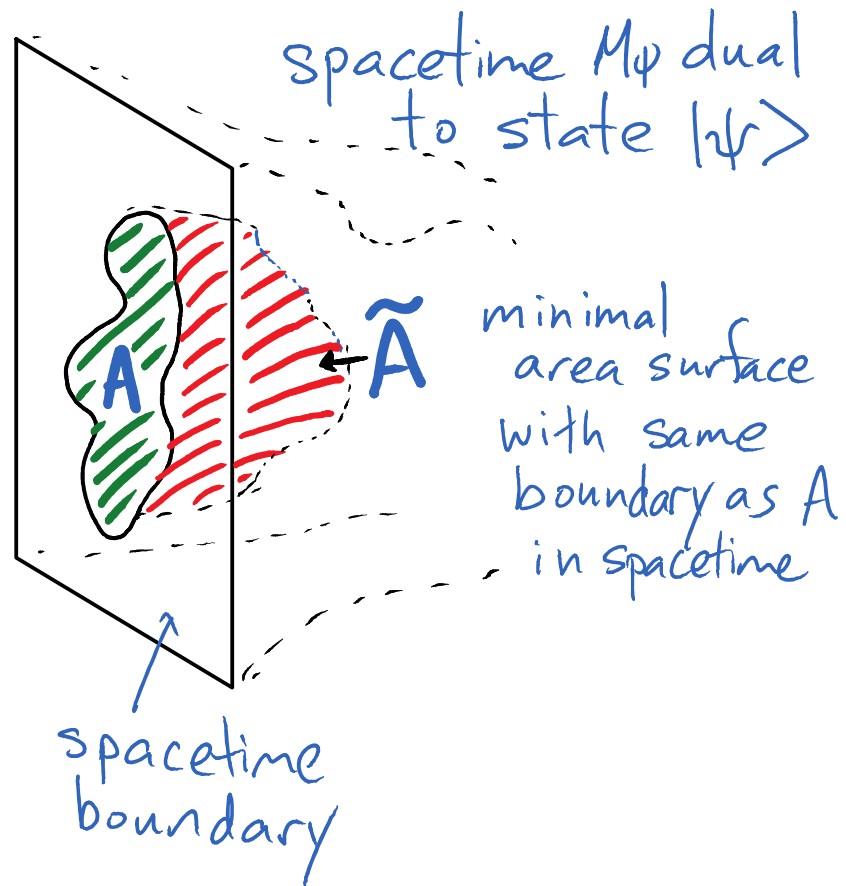
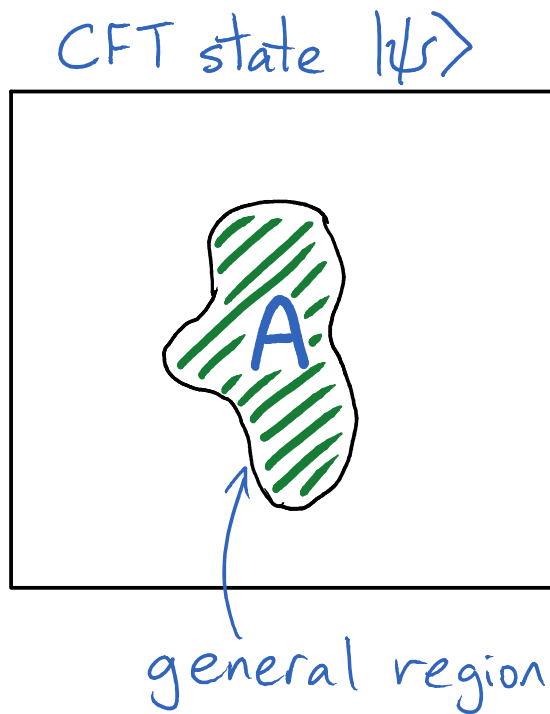


ENTANGLEMENT  
ENTROPY  $S_A$   
(measures how much  
A is entangled w.  
rest of system)



$$= \frac{1}{4G} \times \text{Area of minimal surface } \tilde{A} \text{ in dual spacetime}$$

# GEOMETRY FROM ENTANGLEMENT



$|\psi\rangle \rightarrow$  calculate  $S_A$  for many A  $\rightarrow$  find  $M_\psi$  s.t.  
 $\text{Area}(\tilde{A}) = S_A$

can (plausibly) reconstruct geometry from entanglement!

Past year: gravitational dynamics from entanglement

N. Lashkari, M. McDermott, MVR

T. Faulkner, M. Guica, T. Hartman, R. Myers, MVR

B. Swingle, MVR.

also:

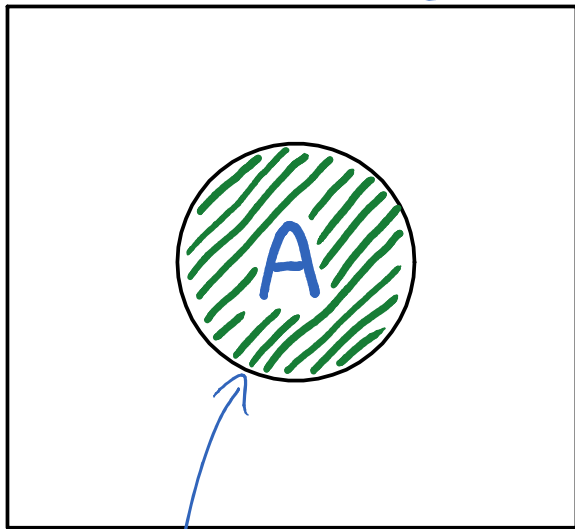
If spacetime geometry is a manifestation of QFT entanglement,  
can we understand spacetime dynamics from entanglement physics?

↑  
Geometries must satisfy  
Einstein's Equations

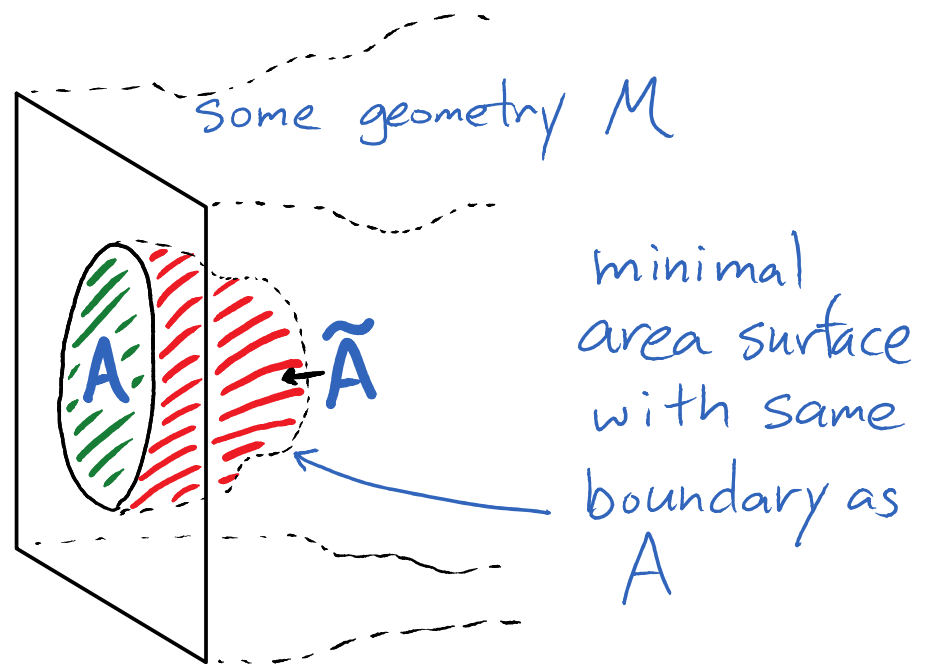
# Will assume:

There exist CFTs for which entanglement entropies for family of states can be computed geometrically.

excited state  $|\psi\rangle$

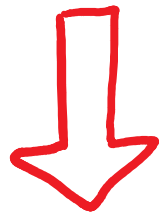


ball shaped region



Which geometries are possible?

Entanglement entropies obey constraints

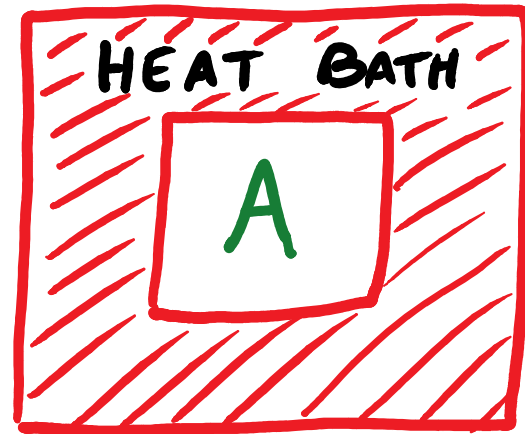


Entropy-area connection

Constraints on geometries

# A QUANTUM FIRST LAW OF THERMODYNAMICS:

Starting in thermal state with temp.  $T$ :



For ANY variation of the state

$$\delta S_A = \frac{1}{T} \delta \langle E \rangle$$

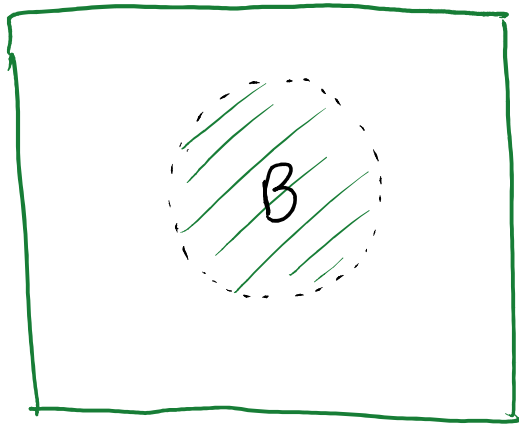
entanglement entropy

energy (Hamiltonian)



Similar "First Law" for perturbations to ground state of a CFT

Blanco, Casini,  
Hung, Myers



$$|0\rangle \rightarrow |0\rangle + \delta|\psi\rangle$$

For any ball B:

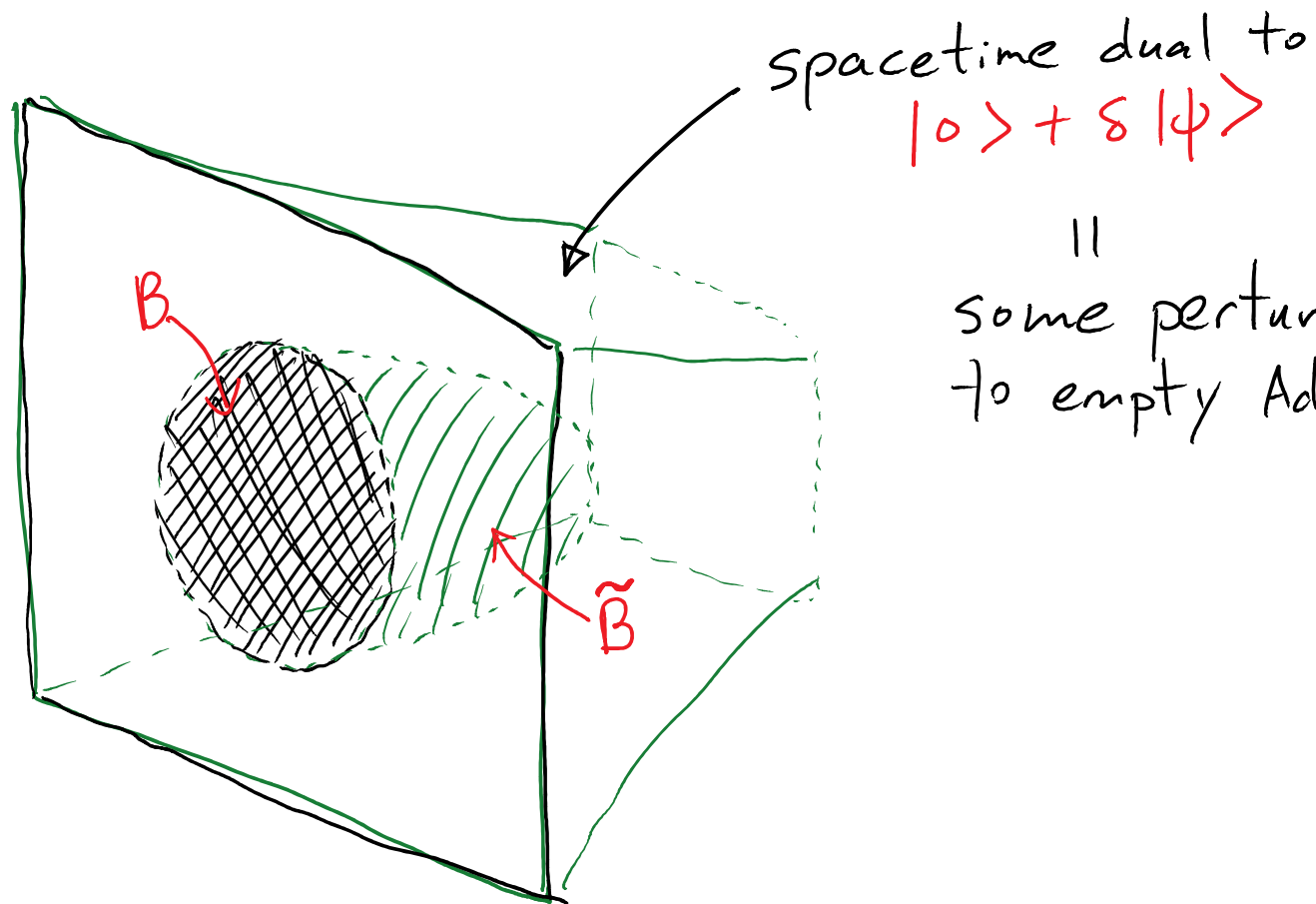
$$\delta S_B = \delta \langle E_B \rangle$$

$$E_B = \int_B dx f(x) \cdot \langle \rho(x) \rangle$$

specific  
function on B

Energy density

Translate to statement about dual spacetime:



||  
some perturbation  
to empty AdS space

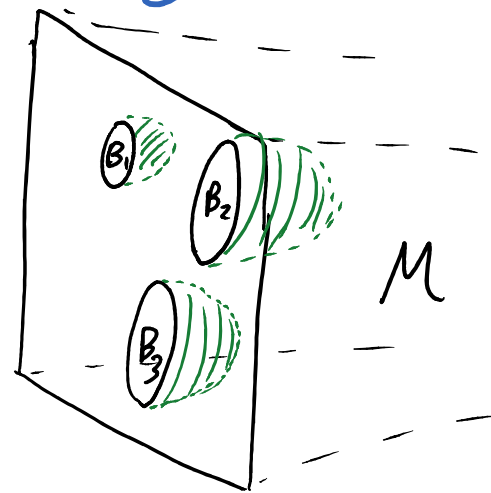
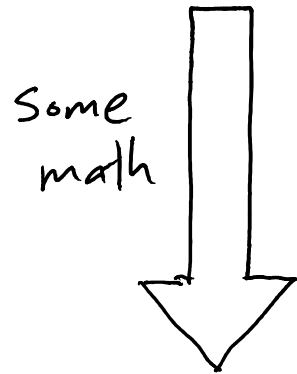
$$dS_B = dE_B$$



relates metric perturbation  
on  $\tilde{B}$  to metric perturbation  
on  $B$  (use Ryu + Takayanagi)

CONSTRAINT ON DUAL SPACETIME

Set of constraints for all ball-shaped regions  $B$



GEOMETRY  $M$  MUST SATISFY EINSTEIN'S EQUATIONS!!  
(linearized about Anti-de-Sitter)

# SUMMARY

structure of entanglement  
in QFT

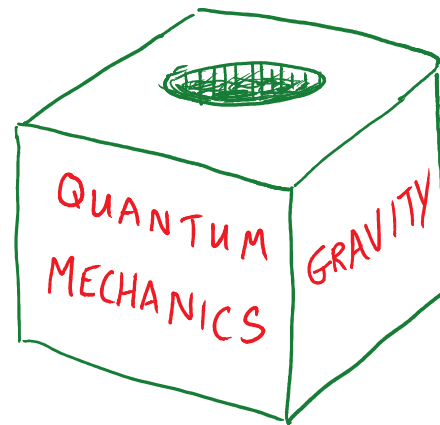
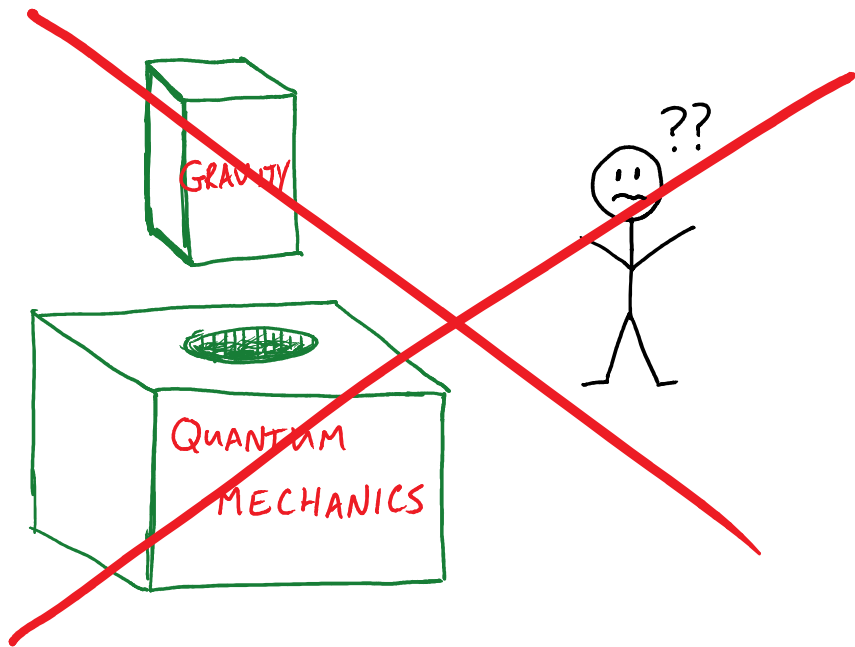


spacetime  
geometry

constraints on QFT  
entanglement



Einstein Equations  
(so far: linearized)



## Further results:

- Including first quantum corrections on gravity side gives  $\langle T_{\mu\nu} \rangle$  as source for Einstein Equations

(linearized Einstein) +  $T_{\mu\nu}$  source + (B.H. entropy = area)

→ Full non-linear Einstein (additional assumptions?)

- Entanglement constraints for finite perturbations imply (averaged) energy conditions