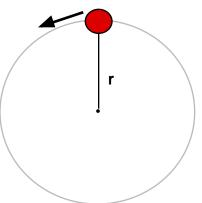
PHYS120 Lecture 23 - Rotational kinematics

Demonstrations: •ball on a string, wheel *Text*: Fishbane 9-2 *Problems*: 10, 13, 57 from Ch. 9

What's important: •definitions of angular variables •angular kinematics

## **Rotational Kinematics**

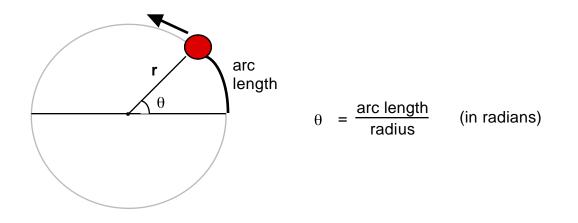
Consider the motion of an object at the end of a string as it moves in a circular path at constant speed,



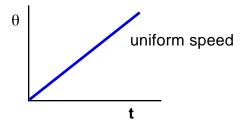
The speed  $|\mathbf{v}|$  and radius *r* are constant, but the position vector **r** and velocity **v** change continuously, having the form of trigonometric functions. Polar coordinates provide a simpler representation of this situation than Cartesian coordinates do, because only one polar coordinate changes with time, namely the angle  $\theta$ . We now develop a description of angular motion based upon polar coordinates, starting with uniform circular motion.

## **Uniform Circular Motion**

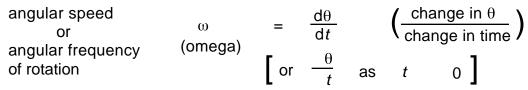
In two dimensions, the position of an object can be described using the polar coordinates r and  $\theta$ . If the object is moving in uniform circular motion, then r is constant, and the time dependence of the motion is contained in  $\theta$ . As a kinematic variable,  $\theta$  is the angular analogue of position.



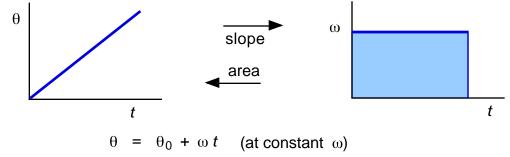
The sign convention is that  $\theta$  increases in a **counter-clockwise** direction. As a function of time,  $\theta$  looks like (for uniform circular motion):



Now, if  $\theta$  is the angular analogue of position, then the slope of the  $\theta$  vs. *t* graph is the angular analogue of velocity. We define



We can go back and forth from  $\theta$  to  $\omega$  by slopes and areas, just like with **x** and **v**. For example, from the area under the  $\omega$  vs. *t* curve, we find



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Two other quantities used to describe uniform circular motion are the period **T** and the frequency *f*. During the period *T*, the object sweeps through one complete revolution, or 2 radians. Hence:

The frequency f is  $T^{-1}$ , so we also have

 $\omega = 2 \quad f$ Note:  $\omega = 1 \quad 1 \text{ radian / second}$   $f = 1 \quad 1 \text{ revolution / second}$ 

## Angular and linear links

We have a relation between *x* and  $\theta$  already, namely

[arc length] =  $r\theta$ .

To obtain a relation between **v** and , we return to the definition of speed:

speed = d[*arc length*] / dt = d( $r\theta$ ) / dt =  $r d\theta$  / dt =  $r\omega$ .

 $v = \omega r$  links linear and angular speeds

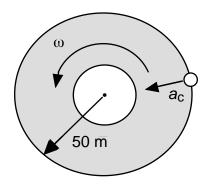
Another linking relation is through the centripetal acceleration  $\mathbf{a}_{c}$ . From previous work, there is a centripetal acceleration even at constant speed:

$$a_{\rm C} = \frac{V^2}{r} = \frac{(\omega r)^2}{r}$$

$$a_{\rm C} = \omega^2 r \qquad \text{(or } a_{\rm C} = \omega v\text{)}$$

## Example

We have a doughnut-shaped space station on which we wish to produce artificial gravity through rotation. What must  $\omega$  be so that  $a_c = g$  at 50 m from the centre of the station?



Using  $a_c = \omega^2 r$ , we find

 $\omega^2 = 9.8 / 50$  -->  $\omega = (9.8/50) = 0.443$  rad/sec.

Note that the units are OK.

Other quantities which we can calculate in this example are:

$$T = \frac{2}{\omega} = \frac{2 \cdot 3.142}{0.443} = 14.2 \text{ sec}$$

which is moderately fast and would not encourage the occupants of the space station to look out the windows. The corresponding speed at the perimeter of the station is:

$$v = \omega r$$
  $v = 0.44_3 \cdot 50 = 22.2 \text{ m/s}$   
= 80 km/hr