

## Lecture 20 - Schrödinger equation in 3D

*What's important:*

- Schrödinger equation in 3D
- angular momentum operators

*Text:* Gasiorowicz, Chap. 10

### 3D Schrödinger equation

As discussed in Lec. 17, the time-independent Schrödinger equation for a free particle in three dimensions reads:

$$-\frac{\hbar^2}{2m} \frac{d^2}{dx^2} + \frac{d^2}{dy^2} + \frac{d^2}{dz^2} u_E(x, y, z) = E u_E(x, y, z). \quad (1)$$

Adding a time-independent potential can easily be accommodated through

$$-\frac{\hbar^2}{2m} \frac{d^2}{dx^2} + \frac{d^2}{dy^2} + \frac{d^2}{dz^2} u_E(x, y, z) + V(x, y, z) u_E(x, y, z) = E u_E(x, y, z) \quad (2)$$

Now, for many situations of interest, the potential:

- is between two objects, both of whose motion must be considered
- depends only on the radial separation  $r$ , so-called central potentials.

Two-particle wavefunctions in one-dimension were introduced in Lec. 16, where we showed that if the potential energy depended only on the relative separation  $x_{\text{rel}}$ , the wavefunction could be written as a product of independent wavefunction for the relative motion and the cm motion. The same is true for three dimensions for central potentials  $V(r)$ . To review the notation (**bold** indicates vectors):

$$\mathbf{R}_{\text{cm}} = (m_1 \mathbf{r}_1 + m_2 \mathbf{r}_2) / (m_1 + m_2)$$

and (3)

$$\mathbf{r} = \mathbf{r}_1 - \mathbf{r}_2.$$

With this replacement, the arguments of the plane wave

$$\mathbf{p}_1 \mathbf{r}_1 + \mathbf{p}_2 \mathbf{r}_2 = \mathbf{P}_{\text{cm}} \mathbf{R}_{\text{cm}} + \mathbf{p} \mathbf{r}, \quad (4)$$

where the total and relative wavevectors are

$$\mathbf{P}_{\text{cm}} = \mathbf{p}_1 + \mathbf{p}_2 \quad \text{and} \quad \mathbf{p} = (m_2 \mathbf{p}_1 - m_1 \mathbf{p}_2) / (m_1 + m_2). \quad (5)$$

The total kinetic energy then becomes

$$E = P_{\text{cm}}^2 / 2M_{\text{total}} + p^2 / 2\mu. \quad (6)$$

Because the kinetic energy separates cleanly into two pieces, and the potential energy depends on relative separation, the wavefunction separates into a plane-wave part describing the cm, and a relative part which satisfies the Schrödinger equation with a reduced mass:

$$U(\mathbf{R}_{\text{cm}}, \mathbf{r}) = \exp(i\mathbf{P}_{\text{cm}}\mathbf{R}_{\text{cm}}/\hbar) \cdot u(\mathbf{r})$$

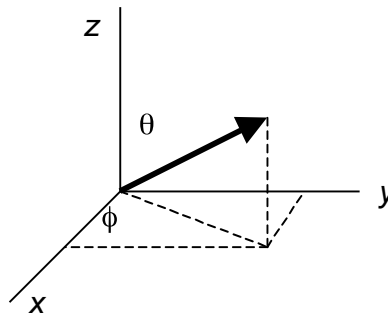
where  $u(\mathbf{r})$  satisfies the time-independent SE:

$$-\frac{\hbar^2}{2\mu} \nabla^2 u(\mathbf{r}) + V(r) u(\mathbf{r}) = E u(\mathbf{r}). \quad (7)$$

Many of the problems of interest in the remaining part of this course are drawn from atomic and molecular physics, where the mass of the two-particle system is concentrated in the nucleus, such that the total mass is close to the nuclear mass and the reduced mass is close to the electron mass. As a result, we will often write Eq. (7) with  $\mu$  replaced by  $m_e$ ; but we should keep in mind that the *exact* solution involves the reduced mass.

### Central potentials

That the potential in Eq. (7) depends on  $r$ , rather than  $\mathbf{r}$ , suggests that it may be useful to represent  $\nabla^2$  in spherical polars, rather than cartesian coordinates. Recalling the definition



It takes repeated application of the chain rule to establish that:

$$\begin{aligned} \nabla^2 &= \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \\ &= \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2}{\partial \phi^2} \end{aligned} \quad (8)$$

The first term involves only  $r$ , giving us some hope that an approach based on the separation of variables might help. As a first step, write the wavefunction  $u(r, \theta, \phi)$  as a product state:

$$u(r, \theta, \phi) = R(r) \cdot Y(\theta, \phi) \quad (9)$$

and substitute into Eq. (7), transposing  $E \cdot u(r, \theta, \phi)$  in so doing:

$$-\frac{\hbar^2}{2m} Y(\theta, \phi) \frac{1}{r^2} \frac{\partial}{\partial r} r^2 \frac{\partial}{\partial r} R(r) + R(r) \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta} + \frac{1}{\sin \theta} \frac{\partial^2}{\partial \phi^2} Y(\theta, \phi) - [E - V(r)] R(r) Y(\theta, \phi) = 0$$

So long as  $V(r)$  has no derivatives, we can divide by  $R(r)$  without worry, leading to (after multiplying by  $r^2$ ):

$$\frac{1}{R} \frac{\partial}{\partial r} r^2 \frac{\partial}{\partial r} R + \frac{2mr^2}{\hbar^2} [E - V(r)] = -\frac{1}{Y \sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta} + \frac{1}{\sin \theta} \frac{\partial^2}{\partial \phi^2} Y \quad (10)$$

To make the expression look less cumbersome, the arguments on  $RY$  have been dropped. This looks promising: the LHS depends only on  $r$ , and the RHS depends only on  $\theta, \phi$ , so both sides must equal a constant, which we define with some foreshadowing as  $\ell(\ell+1)$ . This gives us two equations:

$$\frac{1}{r} \frac{\partial}{\partial r} r^2 \frac{\partial}{\partial r} R + \frac{2mr^2}{\hbar^2} [E - V(r)] R = \ell(\ell+1) R \quad (11)$$

and

$$-\frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta} + \frac{1}{\sin^2 \theta} \frac{\partial^2}{\partial \phi^2} Y = Y \ell(\ell+1) \quad (12)$$

We'll return to the first of these equations, called the *radial equation*, in a few lectures. The solution to the radial equation obviously depends on the choice of  $V(r)$ . But the angular equation is universal - valid for any central potential. We'll solve the  $Y$  functions in the next lecture, but to aid our interpretation, a diversion into the representation of angular momentum is in order.

### Angular momentum in polar coordinates

We have already established an operator representation for the linear momentum  $\mathbf{p}$ . In the following lectures, we interpret the functions  $Y$  in terms of angular momentum  $\mathbf{L}$ , so it is worthwhile to establish at least a few of the angular momentum operator here, as well as its representation in spherical polar coordinates. Now, the classical form of  $\mathbf{L}$  is

$$\mathbf{L} = \mathbf{r} \times \mathbf{p}.$$

When the usual quantum substitution  $p_x \rightarrow -i\hbar \partial/\partial x$  is performed, the order of  $\mathbf{r}$  and  $\mathbf{p}$  is even more important than in the classical cross product. For example,

$$L_z = xp_y - yp_x = -i\hbar \left( x \frac{\partial}{\partial y} - y \frac{\partial}{\partial x} \right) \quad (13)$$

In polar coordinates, this operator has the form

$$L_z = -i \hbar \partial / \partial \phi \quad (14)$$

which we establish by consistency. Recalling the transformation between coordinate systems

$$x = r \sin \theta \cos \phi \quad y = r \sin \theta \sin \phi \quad z = r \cos \theta$$

then

$$\begin{aligned} L_z &= -i \hbar \frac{\partial}{\partial \phi} = -i \hbar \frac{x}{\phi} \frac{\partial}{\partial x} + \frac{y}{\phi} \frac{\partial}{\partial y} + \frac{z}{\phi} \frac{\partial}{\partial z} \\ &= -i \hbar \frac{r \sin \theta \cos \phi}{\phi} \frac{\partial}{\partial x} + \frac{r \sin \theta \sin \phi}{\phi} \frac{\partial}{\partial y} + \frac{r \cos \theta}{\phi} \frac{\partial}{\partial z} \\ &= -i \hbar \frac{-r \sin \theta \sin \phi}{x} + r \sin \theta \cos \phi \frac{1}{y} + 0 \cdot \frac{1}{z} \\ &= -i \hbar \left( -y \frac{\partial}{\partial x} + x \frac{\partial}{\partial y} \right) \end{aligned}$$

The last line is just Eq. (13) again. Similar expressions can be obtained for  $L_x$  and  $L_y$ :

$$L_x = -i \hbar \left( -\sin \phi \frac{\partial}{\partial \theta} - \cot \theta \cos \phi \frac{\partial}{\partial \phi} \right) \quad (15)$$

$$L_y = -i \hbar \left( \cos \phi \frac{\partial}{\partial \theta} - \cot \theta \sin \phi \frac{\partial}{\partial \phi} \right) \quad (16)$$

Eqs. (14) - (16) can be combined to yield:

$$L^2 = -\hbar^2 \left( \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta} + \frac{1}{\sin^2 \theta} \frac{\partial^2}{\partial \phi^2} \right) \quad (17)$$

Eq. (17) has the same form as (13), which is why we have made the identification with  $\ell(\ell+1)$ . Why we chose this instead of  $\ell^2$  will become apparent in the next two lectures.